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QCD radiation in transPlanckian scattering

N. Arkani-Hamed, S. Dimopoulos and G. R. Dvali, Phys. Lett. B **429**, 263 (1998) [arXiv:hep-ph/9803315].

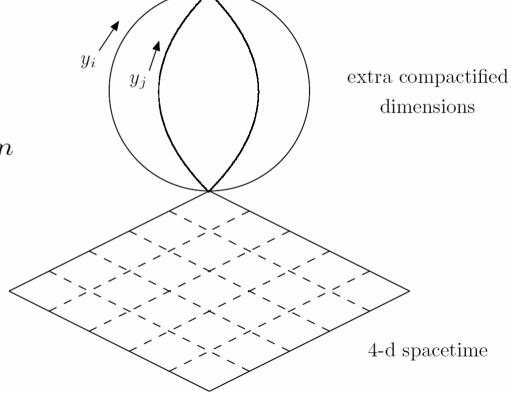


$$S_{\text{bulk}} = -\frac{1}{2} \int d^{4+n} x \sqrt{-g^{(4+n)}} M_{\star}^{n+2} R^{(4+n)}$$

"integrating out" the extra dimensions:

$$M_{\rm Pl}^2 = M_{\star}^{n+2} (2\pi r)^n$$

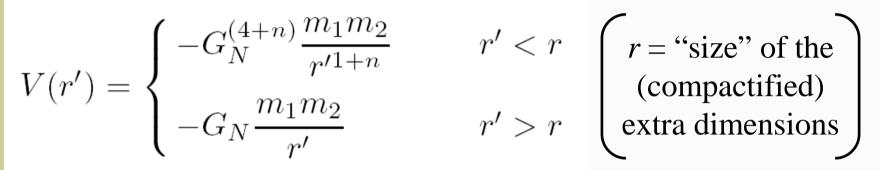
Quantum gravity scale could be much lower than MPI !



Fundamental ASSUMPTION:

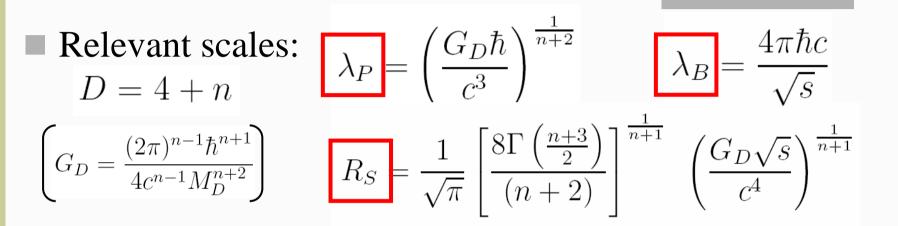
only gravity propagates in the extra dimensions

Expected deviations from Newtonian gravity:



Experimental bound: $r \le 0.2 \text{ mm}$ For example: $M_* \sim 10^3 \text{ GeV}$ $r \sim 2 \cdot 10^{-17} 10^{\frac{32}{n}} \text{ cm}$ $M_* \sim 10^3 \text{ GeV}$ and $n \ge 2$ is a serious possibility!

2) Transplanckian scattering

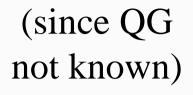


Peculiar feature: NEW SCALE

- $\left(\frac{G_D S}{\hbar 5}\right)^{\frac{1}{n}}$ Cannot be defined if n=0Goes to infinity if $\hbar \to 0$ w
 - Goes to infinity if $\hbar \rightarrow 0$ with G_D constant

"Size of classical region" (assume $b_c \gg R_S$)

Model independent features



- Impact parameter $b \gg R_S \rightarrow$ weak gravitational field (linearization)
 - $\sqrt{s} \gg M_D \qquad \Rightarrow \qquad R_S \gg \lambda_P \gg \lambda_B$

 \rightarrow quantum-gravity effects are small

(we can use just QM and linearized GR)

Forward scattering at small angles \rightarrow eikonal regime, predictive computation is possible

TransPlanckian eikonal regime:

$$M_D/\sqrt{s} \ll 1, \qquad -t/s \ll 1$$

REMARKABLE FACTS:

- Predictivity
- Model independence

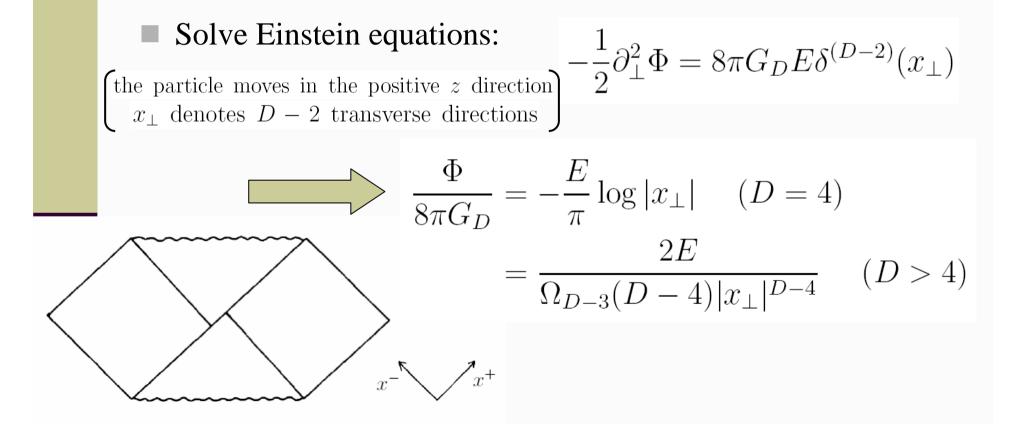
D. Kabat and M. Ortiz, Nucl. Phys. B 388, 570 (1992) **FIRST APPROACH:** Eikonalization evaluate leading behaviour at high energy and small angles by summing an infinite number of diagrams - Take on shell vertices approximation: -Use: $\frac{1}{(p+k)^2 + m^2 - i\epsilon} \approx \frac{1}{2p \cdot k - i\epsilon}$ (resummation) $\mathcal{A}_{\text{eik}} = -2is \int d^2 b_{\perp} e^{iq_{\perp}b_{\perp}} \left(e^{i\chi} - 1\right) \quad t = q^2 \simeq -q_{\perp}^2$ $\chi(b_{\perp}) = \frac{1}{2s} \int \frac{d^2 q_{\perp}}{(2\pi)^2} e^{-iq_{\perp}b_{\perp}} \mathcal{A}_{\text{Born}}(q_{\perp}^2) = \left(\frac{b_c}{b}\right)^n \qquad \left\{ \begin{array}{l} -\text{ No UV sensitivity} \\ -\text{ Spin independent} \end{array} \right.$

G. 't Hooft, Phys. Lett. B198 (1987) 61 P.C. Aichelburg and R.U. Sexl, Gen. Rel. Grav. 2 (1971) 303

SECOND APPROACH: Shock wave (generalized)

Replace one particle by its AS shock wave:

$$ds^{2} = -dx^{+}dx^{-} + \Phi(x_{\perp})\delta(x^{-})(dx^{-})^{2} + dx_{\perp}^{2}$$

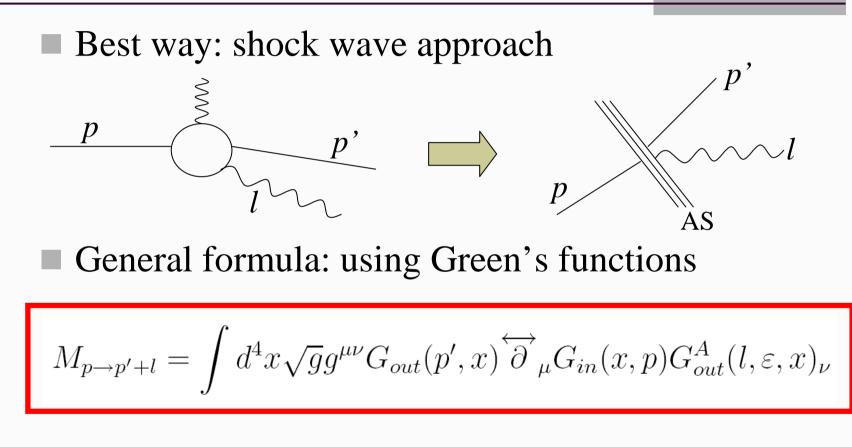


$$\left(\begin{array}{cc} \text{AS shock} \\ \text{wave:} \end{array} & ds^2 = -dx^+ dx^- + \Phi(x_\perp)\delta(x^-)(dx^-)^2 + dx_\perp^2 \end{array} \right)$$

$$\left(\begin{array}{cc} \text{Perform a (discontinuous) change of coordinates so that the metric becomes continuous across } x^- = 0 & : \\ & \left\{ \begin{array}{c} x^+ = x'^+ + \theta(x^-)[\Phi(x_\perp) + \frac{(\partial_\perp \Phi(x_\perp))^2}{4}x^-] \\ x^- = x'^- \\ x^i = x'^i + \theta(x^-)\frac{\partial_i \Phi(x_\perp)}{2}x^- & (i = 1, 2) \end{array} \right) \right)$$

$$\left(\begin{array}{c} \text{Solve Klein-Gordon equation in these new coordinates (imposing continuity of the wavefunction):} \\ \text{before collision:} & \psi(x) = \exp(+ip'x) = \exp(-iE'(z+t)) \\ \text{after collision:} & \psi(x') = \exp(-iE'x^{+\prime}) = \exp(-iE'(x^+ - \Phi(x_\perp))) \\ \text{eikonal amplitude:} \\ \hline e^{i\chi} = \exp(iE'\Phi) \\ \end{array} \right)$$

3) PROJECT: Including QCD radiation



Main point: evaluation of this integral! (Green's functions are obtained via 't Hooft method)

STUDY OF M: leading log corrections

Including all factors: total cross section = without gluon
 + large logarithms (IR singularity if quark massless)

 $\rightarrow AP equations$ (the usual DIS scaling violation)

redefine scale dependent parton distribution functions so that main corrections are reabsorbed

• What we find:

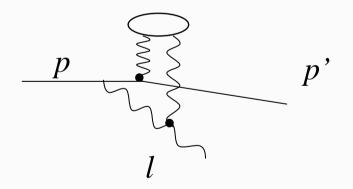
 $\begin{array}{c} q_{\perp}b_c \lesssim 1 \\ q_{\perp}b_c \gg 1 \end{array} \qquad \square \end{array}$

scale =
$$q_{\perp}$$
 as usual (JJ)
different
scale! $(q_{\perp}b_c)^{\frac{1}{n+1}}$

Really a very peculiar fact! Physical interpretation: saddle point

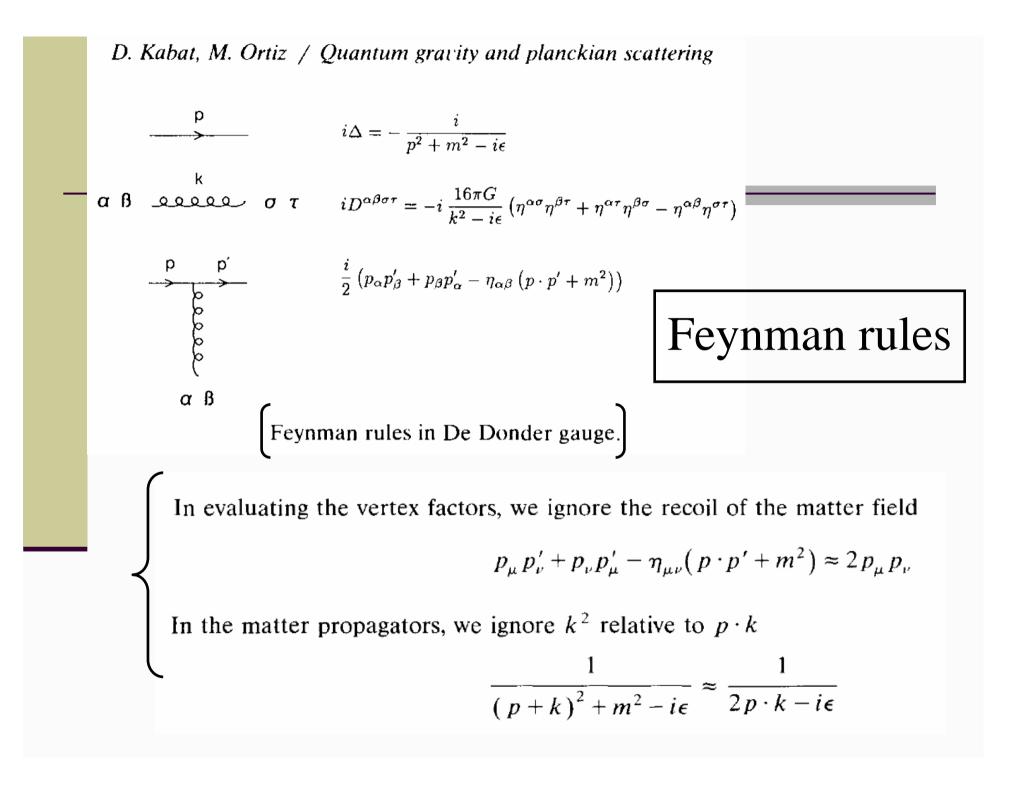
4) Conclusions

- Generalization of 't Hooft shock wave method in transPlanckian scattering in order to include radiation
- For large momentum transfer $(q_{\perp}b_c \gg 1)$ the PDF must be normalized at a scale different than usual. Physical reason: RESCATTERING EFFECTS

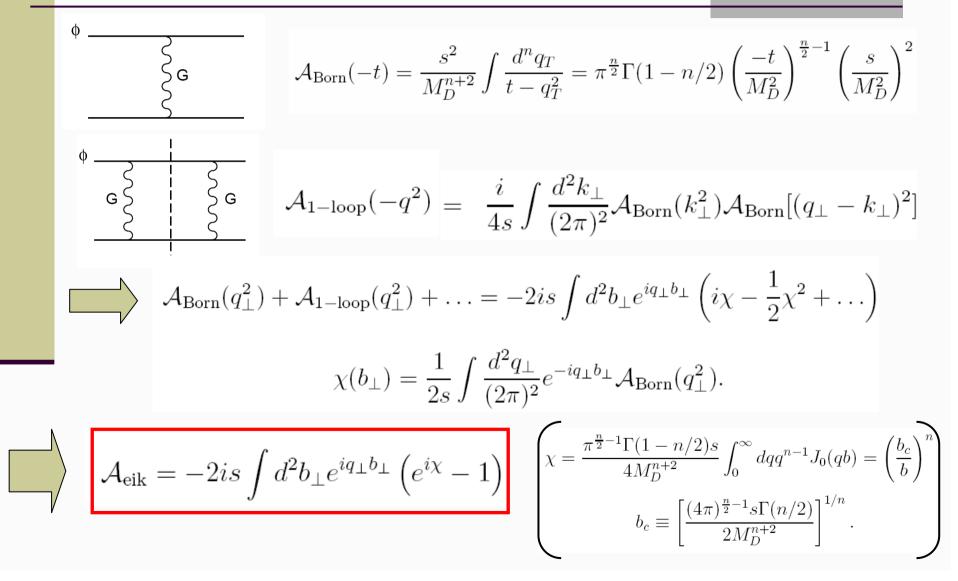


Future developments: WORK IN PROGRESS

Extra Slides



Eikonalization vs Shock wave: eikonal resummation



Eikonalization vs shock wave: shock wave computation:

Let us consider the plane wavefunction colliding with the classical AS shock wave. The wavefunction describes a particle of energy E' moving in the negative z direction. Thus p' = (E', -E', 0) and before the collision the wavefunction is:

$$\psi(x) = \exp(+ip'x) = \exp(-iE'(z+t))$$

After the collision the wavefunction remains continuous in the x' coordinates, i.e.

$$\psi(x') = \exp(-iE'x^{+\prime}) = \exp(-iE'(x^{+} - \Phi(x_{\perp}))) , \quad x^{-} = \varepsilon > 0 \text{ small}$$

(i.e. immediately after collision). Thus we get eikonal amplitude:

$$e^{i\chi} = \exp(iE'\Phi) \qquad \qquad \left(\Omega_k = \frac{2\pi^{(k+1)/2}}{\Gamma((k+1)/2)} \right)$$

Check:

$$\begin{split} \chi &= 8\pi G_D \frac{2}{\Omega_{D-3}(D-4)b^n} EE' = \frac{(2\pi)^n}{M_D^{2+n}} \frac{\Gamma((k+1)/2)}{2\pi^{(k+1)/2}}|_{k=n+1} \frac{2}{n} \frac{s}{4b^n} \\ &= \frac{(2\pi)^n}{M_D^{2+n}} \frac{\Gamma(\frac{n}{2}+1)}{2\pi^{\frac{n}{2}+1}} \frac{2}{n} \frac{s}{4b^n} = \frac{2^{n-2}\pi^{\frac{n}{2}-1}}{M_D^{2+n}} \frac{\Gamma(\frac{n}{2})}{2} \frac{s}{b^n} = \frac{(4\pi)^{\frac{n}{2}-1}}{M_D^{2+n}} \frac{\Gamma(\frac{n}{2})}{2} \frac{s}{b^n} = \frac{1}{2\pi^{\frac{n}{2}+1}} \frac{1}{2\pi^{\frac{n}{2}+1}}$$

Example of signal at the LHC

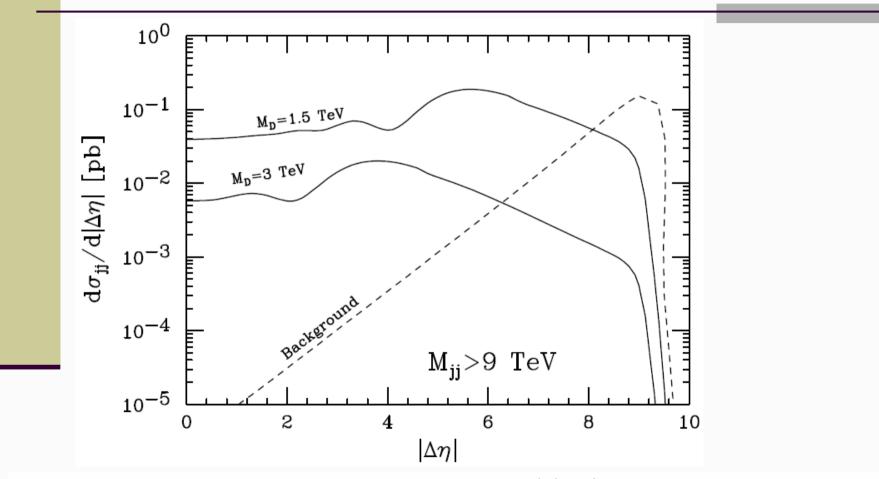


Figure 4: The di-jet differential cross section $d\sigma_{jj}/d|\Delta\eta|$ from eikonal gravity for n = 6, $M_{jj} > 9$ TeV, when both jets have $|\eta| < 5$ and $p_T > 100$ GeV, and for $M_D = 1.5$ TeV and 3 TeV. The dashed line is the expected rate from QCD.

Green's functions

0

•
$$G_{out}(p',x) = e^{-ip'x}$$
, $x^- > 0$
• $G_{out}^A(l,\varepsilon,x)_\mu = \varepsilon_\mu e^{-ilx}$, $x^- > 0$, $\varepsilon l = 0$
• $G_{in}(p',x) = e^{+ipx}$, $x^- < 0$
• $G_{in}(p,x) = e^{-i\left\{\frac{1}{2}p^{-}[x^+ - \Phi(x^i)] - p^i x^i\right\}}$, $x^- = 0^+$
 $= \int \frac{d^2k^i}{(2\pi)^2} e^{-i\left(\frac{(p^i + k^i)^2}{2p^-}x^- + \frac{p^-}{2}x^+ - (p^i + k^i)x^i\right)} A(k^i, p^-)$ $(x^- > 0)$
• $G_{out}(p',x) = e^{i\left\{\frac{1}{2}p'^-[x^+ + \Phi(x^i)] - p'^i x^i\right\}}$, $x^- = 0^-$
 $= \int \frac{d^2k^i}{(2\pi)^2} e^{i\left(\frac{(p'^i - k^i)^2}{2p'-}x^- + \frac{p'^-}{2}x^+ - (p'^i - k^i)x^i\right)} A(k^i, p'^-)$ $(x^- < 0)$

$$\begin{aligned} G_{+}^{A} &= \int \frac{d^{2}k^{i}}{(2\pi)^{2}} e^{-i\tilde{l}x} A(k^{i}, l^{-})\varepsilon_{+} \qquad G_{-}^{A} = \int \frac{d^{2}k^{i}}{(2\pi)^{2}} e^{-i\tilde{l}x} A(k^{i}, l^{-})[\varepsilon_{-} + \frac{(k^{i})^{2}}{(l^{-})^{2}}\varepsilon_{+} + \frac{\varepsilon_{i}k^{i}}{l^{-}}] \\ G_{i}^{A} &= \int \frac{d^{2}k^{i}}{(2\pi)^{2}} e^{-i\tilde{l}x} A(k^{i}, l^{-})[\varepsilon_{i} + \frac{2k^{i}}{l^{-}}\varepsilon_{+}] \\ &\qquad (\tilde{l}^{-}, \tilde{l}^{+}, \tilde{l}^{i}) = (l^{-}, \frac{(l^{i} - k^{i})^{2}}{l^{-}}, l^{i} - k^{i}) \end{aligned}$$

AP equations (roughly, many factors missing)

$$\begin{aligned} \frac{d\sigma}{d^2 q_{\perp} dx} &= \delta(x-1) \frac{1}{4\pi^2} \left| A_{reg}(q_{\perp};p^-) \right|^2 \\ &+ \frac{1}{(2\pi)^5} \delta(x-x(z)) \frac{z \, dz}{1-z} \int d^2 l_{\perp} \left| A_i(q_{\perp},l_{\perp},z;p^-) \right|^2 \\ &\propto \left\{ \frac{1}{(n+1)} \log(b_c q) - \log(b_c \mu) \right\} |A_{reg}(q,zp^-)|^2 \\ &+ \left\{ \frac{1}{(n+1)} \log(b_c q) - \log(b_c \mu) \right\} |A_{reg}(q,(1-z)p^-)|^2 + \cdots \end{aligned}$$

 $(\leftrightarrow \text{new scale})$